REPRINT AND CIRCULAR SERIES OF THE NATIONAL RESEARCH COUNCIL

INDEXING OF SCIENTIFIC ARTICLES

By Gordon S. Fulcher Corning Glass Works, Corning, New York

Including an Analytic Index of the Astrophysical Journal, Vol. 51-54, 1920-1921

Announcement Concerning Publications of the

National Research Council

The Proceedings of the National Academy of Sciences

has been designated as the official organ of the National Research Council for the publication of accounts of research, committee and other reports. and minutes.

Subscription rate for the "Proceedings" is \$5.00 per year. Business address: Home Secretary, National Academy of Sciences, Smithsonian Institution, Washington, D.C.

The Bulletin of the National Research Council

presents contributions from the National Research Council, other than proceedings, for which hitherto no appropriate agencies of publication have existed.

The "Bulletin" is published at irregular intervals. The subscription price, postpaid, is \$5.00 per volume of approximately 500 pages. Numbers of the "Bulletin" are sold separately at prices based upon the cost of manufacture.

The Reprint and Circular Series of the National Research Council

renders available for purchase, at prices dependent upon the cost of manufacture, papers published or printed by or for the National Research Council.

Orders for the "Bulletin" or the "Reprints and Circulars" of the National Research Council, accompanied by remittance, should be addressed: Publication Office, National Research Council, 1701 Massachusetts Avenue, Washington, D.C.

INDEXING OF SCIENTIFIC JOURNALS

By GORDON S. FULCHER

The importance of the service which the subject-indexes of journals may render to scientific research is not generally realized. Yet, in order that the best methods may be chosen and unnecessary duplication avoided, each research should, of course, be based on as complete a knowledge of past results as may be obtained; and the task of guiding the scientist to those parts of the enormous accumulation of scientific literature which relate to his work naturally devolves largely upon the subject-indexes. Also, since the greater the amount of time and effort required by scientists to secure the information needed for effective research, the greater will be the proportion discouraged from attempting it and the greater will be the extent to which the research done will be either less efficient or more delayed than it should be, therefore the responsibility resting upon these indexes, particularly the indexes of abstracting and listing journals, to perform this important task well, is very great.

The question then arises, What characteristics must the indexes have if they are to give the maximum service?

The indexes now provided are of various types and degrees of usefulness. The simplest is a mere *unclassified index of titles*, which are arranged alphabetically by their significant words. The indexes of the *Astrophysical Journal*, the *Physical Review*, and many other journals are of this type.

An improved type is the *classified index of titles* in which the titles are arranged under a limited number of subject headings so as to bring related titles together, the classification being based, however, wholly on the titles. The indexes of *Science Abstracts*, *Journal de Physique*, and others are of this type.

The inadequacy of both these types is obvious for it is well known that in most instances a title cannot sufficiently describe all the subjects treated in the article. A paper on "The Atomic Weight of Iodine" may contain results relating to I_2O_5 and to the occlusion of oxygen by glass; and one on "The Flora of Formosa" may describe new species and perhaps new genera.

Ι

A much more serviceable type is the *index of titles classified by content of the articles*, which, while retaining the titles, recognizes their shortcomings and classifies each with more or less precision under headings determined by an examination of the article itself or an abstract. Thus "The Flora of Formosa" would be listed under the genera, species, and other subjects concerning which new information is given. The *International Catalogue of Scientific Literature* and the cumulated indexes of the H. W. Wilson Company—Agricultural Index, Industrial Arts Index, etc.—are of this character. The value of this type of index is enhanced if, as in the case of the Engineering Index and the card indexes of the Concilium Bibliographicum, the classified titles are supplemented by brief statements as to the contents.

Since, however, most titles are incomplete and many inaccurate, why not disregard them altogether? This is done by the *content index*. In preparing it, the precise subjects dealt with in each article are determined by an analysis of the original or of a reliable abstract, index entries which adequately describe these contents are formulated, and finally these entries are indexed under the proper subject-headings, alphabetically arranged. The index of *Chemical Abstracts* is of this type, being based on an analysis of the abstracts, and is not only one of the largest but also one of the most generally useful indexes now being issued. For the most part, however, the classification is carried through one stage only, with the result that if information on a certain phase of one of the larger subjects is desired, many references may have to be looked up. Moreover, the index depends on abstracts which vary considerably in their standards.

The final step in the development of the subject-index is to base the entries directly—or indirectly through abstracts—on a thorough analysis of the original articles, and to carry the subclassification of the entries through two or three stages so as to make each subdivision complete. Since the articles in the *Astrophysical Journal* for 1920 and 1921 were carefully analyzed in preparing the analytic abstracts¹ which precede the articles, it was possible to prepare

¹ Gordon S. Fulcher, "Scientific abstracting," Science, 54, 291-95, 1921.

from them a very thoroughly classified analytic index. This is reprinted at the end of the paper.

To show the characteristics of the analytic index as compared with those of two types of title index, the entries used by each type to index the same four articles, all relating to photography (Astrophysical Journal, 52, 86; 52, 98; 52, 201; and 53, 349), are collected here:

Index of the Astrophysical Journal (Unclassified Title Index)

Contraction and Distortion on Photographic Plates, Image Image Contraction and Distortion on Photographic Plates Images, Mutual Action of Adjacent Photographic Photographic Images, Mutual Action of Adjacent Photographic Plates, Image Contraction and Distortion on Photographic Sharpness and Resolving Power Photometry and the Purkinje Effect, Photographic Purkinje effect, Photographic Photometry and the Resolving Power, Photographic Sharpness and

Index of Science Abstracts (Classified Title Index)

Photography |

Images, Mutual Action of Adjacent Photographic¹
Photometry, Photographic, and Purkinje-Effect
Plates, Photographic, Image Contraction and Distortion on
Resolving Power, Photographic Sharpness and¹
Vision

Purkinje-effect, and Photographic Photometry

Analytic Index²

Photographic plates; properties contrast functions drying, time of grain size images contraction and distortion mutual action sections

- ¹ These references are put in by analogy with the others, as the 1921 index has not yet been issued.
- ² Headings and subheadings alone are given; see the complete index herewith reprinted for details.

```
intensification
  penetration of light
  resolving power
  sharpness
  shifts of spectrum lines and star images
  theoretical relations
  turbidity
Photometry, photographic, heterochromatic
Purkinje effect, photographic
Spectra
  measurements
    errors possible
  shifts of lines
    photographic
Stars
  measurements from photographs
    errors possible
Sun
  spectrum
    shifts of lines
      photographic
```

The precision and completeness of the analytic index as compared with the title indexes is evident. Moreover, since all the results relating to each subject are indexed together, the analytic index also provides a summary, in outline, of the work done in each field; that is, the reader interested in a particular subject, such as clusters, nebulae, spectra, variables, can determine in what phases of the subject progress has been made by consulting the proper sections of the index. On the other hand, the danger of indexing by titles alone is illustrated by the classification, in the *Index of Science Abstracts*, of the reference to the photographic Purkinje effect under "Vision" whereas the article contains, in fact, no reference to the eye, the indexer being misled by the title.

It may be objected that, while the analytic and other good content indexes are clearly more useful than any title indexes, they are too long and too difficult and laborious to prepare. But a comparison of the analytic index, here reprinted, with the corresponding four separate volume indexes of the *Astrophysical Journal* shows that the analytic index has only twice as many words as the

unclassified title indexes, though containing three times as many references, which are, moreover, precisely and thoroughly classified. Also, the preparation does not require remarkable ability nor should it take an undue amount of the editor's time if the work is well organized. After deciding as to the classification to be adopted, the editor's task is merely to indicate on the abstracts the titles to be used to index the articles (heads and subheads being designated in some convenient way) and later to go carefully through the entries which have been typewritten on slips and arranged alphabetically. The clerical work can be greatly simplified by the use of methods developed by such indexing agencies as *Chemical Abstracts*, the H. W. Wilson Company, and the *New York Times*.

The subject-indexes now serving the various sciences vary widely in type and usefulness, as stated above. Some are very poor, some very good; but none is as complete and precise as it should be if it is to give maximum service to research in its field. It is only a question of time, however, when the workers in each science will come to realize the incalculable value of an efficient abstract journal combined with a complete, precise, and thoroughly classified subject-index, and they will arrange to make the small sacrifices of time and money necessary to secure such uniquely important tools.

ANALYTIC INDEX OF ASTROPHYSICAL JOURNAL FOR 1920 AND 1921, VOLUMES 51–54, BASED ON THE ANALYTIC ABSTRACTS

For references to spectra and spectroscopy of elements and compounds see Spectra.

Absorption of light

by atmosphere, 315–290 μμ; coefficients, **54**, 297 by vapors in King's electric furnace; Kirchhoff's law applied to, **51**, 13

Aethylamine, preparation of; method, 52, 129
Angles, extremely small; interferometer method of measurement, 51, 257

Arc

low-current; anode fall, cathode fall, potential and potential gradient as functions of current, to r amp., for various cathode materials (Ag, C, Cd, Cu, Fe, Sn, Zn, and salts of Ba, Ca, Cs, Na, and Sr), 54, 273

Atmosphere (see also Sky)

absorption of light, 315-290 μμ; coefficients, 54, 297

absorption spectrum (Fraunhofer lines) (see Sun, spectrum)

dispersion of light; measurement; stellar interferometer method suggested, 51, 263 ozone in; amount, daily variations, location, and

suggested origin, 54, 297

Atomic theory, Bohr (see Spectra, theory)

Benzene; dispersion, rotary magnetic, 436-620 μμ; 54, 45

Benzene, nitro-; dispersion, rotary magnetic, radiation pressure between components, theory, 436-620 μμ, **54,** 45 o Sagittae, radial velocity variations, 51, 252 Binaries, spectroscopic (see also Variables) spectrum of Boss 2830, comp. a Geminorum, Boss 373, elements of both orbits, parallax, and W Serpentis; notes, 53, 13 and proper motion, 53, 201 theory (see dynamics, origin, and radiation Boss 3644, Virginis, elements of orbit, 54, 226 pressure) Boss 5026, elements of both orbits, 51, 187 21 π Ursae Minoris, radial velocity, 52, 198 Boss 5591, elements of both orbits, parallax, Binaries, visual (Double stars) and proper motion, 53, 201 13 Ceti, orbit, 52, 110 (see Triple systems) +65°369 Camelopard, radial velocity, 52, 198 a Hercules and comp.; probably optical pair, a Capella, elements determined by interfero-53, 201 meter, 51, 263 measurement of separation, photographic; 40 δ Capricorni, elements of orbit, 54, 127 possible errors due to mutual action, 53, 349 12, 15 K, 42, and 50 Cassiopeiae, radial velocities, Brightness (see Clusters, Comet, Galaxy, Magni-52, 198 tudes, Nebulae, Novae, and Sky) 19 Cephei, radial velocity variations, 51, 252 Brightness ratio, of palladium point to gold point, 13 Ceti, binary component; elements of orbit, 51, 244 52, 110 Class Oes: 10 Cephei, A Cygni, o Sagittae, and Camphor 4 more; radial velocity variations, 51, 252 dispersion, rotary; magnetic and natural; solu-A Cygni, radial velocity variations, 51, 252 tion in ethyl alcohol, 54, 116 refractive indices for 1:1 solution in ethyl X Cygni (Cepheid), spectroscopic orbit, 53, 95 65 τ Cygni, binary component, elements of alcohol; 436-620 μμ, **54**, 116 Carbon disulfide; dispersion, rotary, magnetic; orbit, 53, 144 34 ψ2, 68, 71, +62°1637 Draconis, radial veloci-436-620 μμ, 54, 45 ties, 52, 198 Cathode, for vacuum tube 205 Draconis, elements of both orbits, parallax carbon, limed; preparation, 53, 323 and proper motion, 53, 201 use as source of large current, 53, 323 dynamics; tides on sphere due to second sphere Cathode rays; excitation of light in air; intensity rotating around it, 51, 309 for 1500-3500 volts, 52, 278 comp. a Hercules, elements of orbit, parallax and Cepheids (see Variables) proper motion, 53, 201 Cluster stars Lalande 29330 and 46867, elements of orbit, catalogue, photometric (colors and magnitudes) parallax and proper motion, 53, 201 in Messier 3, 848 stars, 51, 140 magnitudes, absolute (see parallax) in Messier 68, 56 giants, 51, 49 measurement of relative brightness and position colors (see catalogue) distribution (number) (see M3 and M68) of components; interferometer method, relation to magnitude (see M3 and M68) 51, 257, 263 orbits; elements of, including periods, radial variation with radial distance (see M3) velocity curves, mass functions giants (see M68) color; relation to magnitude, general con-Boss 373, 3644, 5026, 5591; a Capella, 49 δ Capricorni, 13 Ceti; X and 65 τ Cygni, clusions, 51, 49 205 Draconis, comp. a Hercules, Lalande magnitudes (see catalogue) 29330 and 46867, OΣ 82) distribution (number) (see M3 and M68) measurement of; interferometer method, 51, relation to color (see M3 and M68) variation with radial distance (see M3) origin, theory of, from nebulae, 51, 309 in Messier 3; photometric analysis, 51, 140: π4 Orionis, photometric study, 51, 218 colors and magnitudes OΣ 82, elements of orbit, parallax, and proper catalogue of 848 stars distribution (number of each color and motion, 53, 201 parallax and absolute magnitude (see Boss 373, magnitude) 5591; 205 Draconis, comp. a Hercules, distribution in space, variation with dis-Lalande 29330 and 46867, OS 82, 53, 201) tance from center relation between color and magnitude proper motion (see: Boss 373, 5591; 205 Draposition co-ordinates of 370 stars conis, comp. a Hercules, Lalande 29330 and 46867, OE 82, 53, 201) variables, 17 probable

radial velocity (see orbits, also +65°369, Camelopard; 12, 15K, 42, and 50 Cassio-

Draconis, 21 m Ursae Minoris)

peiae, Class Oe5; $34 \psi^2$, 68, 71, $+62^{\circ}1637$

in Messier 53 (N.G.C. 5024); variables

positions of 23 shown in photograph, 52, 73 in Messier 56 (N.G.C. 6779); variables

positions of 3 shown in photograph, 52, 73

Cluster stars (cont.) Comets in Messier 68 (N.G.C. 4590); photometric brightness, photographs, behavior of tail, 51, 103 analysis, 51, 49: Morehouse's; rejection of tail, 51, 103 giants, colors, and magnitudes tail, rejection of; instances and stages, 51, 103 catalogue of 56 Constant stars (see Stars) distribution (number of each color and Córdoba Observatory Catalogue magnitude) correction for position of star 12431, 51, 4 relation between color and magnitude Corona (see Sun) variables; colors, magnitudes, positions, and ranges of variation of 28 Dark markings (see Nebulae, dark) in Messier 72 (N.G.C. 6891) Diameters, of planetoids, satellites and stars comparison stars, magnitudes of 29, 52, 232 measurement (see Interferometer, stellar) variables; periods and light curves of 26 Discharge through gases Cepheids, 52, 232 occlusion of RaEm, 54, 285 position of 34, in photograph, 52, 73 Dispersion in Messier 75 (N.G.C. 6864); variables by atmosphere; measurement; interferometer position of 16, in photograph, 52, 73 method suggested, 51, 263 in N.G.C. 7006 and 7789; number, 52, 73 theory, electron, modified to include variation number, in N.G.C. 7006 and 7789, 52, 73 with temperature, 51, 223; note as to number of each color and magnitude (see priority, 53, 326 $M_3, M68)$ Dispersion, rotary (see Camphor, Limonene, variables (see M3, 53, 56, 68, 72, 75; Variables) Sugar, Tartaric acid, 54, 116) Clusters (see also Cluster stars) Dispersion, rotary, magnetic (see Benzene, analysis, photometric (colors and magnitudes of nitro-Benzene, Camphor, Carbon disulfide, the stars) (see Messier 3 and 68) Ethyl iodide, Limonene, a-monobrombrightness of average cluster, 52, 73 Naphthalene, Sugar, Tartaric acid) diameter, of N.G.C. 7006 and 7789, 52, 73 theory, electron, extended form of Messier 68, 71, 49 isotropic transparent media, 54, 45 magnitude, absolute (see parallax) optically active media, 54, 116 Messier 3; photometric analysis, 51, 140 Double stars (see Binaries, visual) Messier 8 (N.G.C. 6530), photograph, 51, 4 Dynamics Messier 16 (N.G.C. 6611), photograph, 51, 4 binary system (see Binaries, spectroscopic) Messier 22 (N.G.C. 6656), photograph, 51, 4 single mass, rotating; equilibrium, 51, 309 Messier 53 (N.G.C. 5024), photograph, 52, 73 potential of distorted ellipsoid, 51, 309 Messier 56 (N.G.C. 6779), photograph, 52, 73 Messier 68 (N.G.C. 4590); analysis, photometric; Eclipse (see Sun) form; computed parallax; photograph, Einstein effect (see Gravitation) 51, 49 Electron theory (see Dispersion and Dispersion, Messier 72 (N.G.C. 6981); parallax, 52, 232; rotary, magnetic) photograph, 52, 73 Electrons; e/m for those active in magnetic Messier 75 (N.G.C. 6864); photograph, 52, 73 rotary dispersion, computed, 54, 45 N.G.C. 7006; diameter, parallax and star Emission of light, in discharge tube intensity as function of energy of exciting counts, 52, 73 N.G.C. 7492, photograph, 52, 73 cathode rays, 52, 278 N.G.C. 7780; diameter; number of stars, Equilibrium (see Dynamics) Ethyl iodide; dispersion, magnetic rotary, 436parallax and absolute magnitude (see Messier 620 μμ, 54, 45 68, 72; N.G.C. 7006) Exploded wire, source of high temperature spectra average cluster, brightness, 52, 73 (see Spectra) computation, methods, 51, 40 appearance and mechanical effects (Plate), 51, 37 globular clusters, forty, 52, 73 Fluorine; preparation of gas by electrolysis, 54, 133 photographs, large scale (60-inch) Messier 53, 56, 72, 75 and N.G.C. 7492, black body, double platinum wound, 51, 244 showing variable stars, 52, 73 electric (see Spectra) Messier 68, showing variables, 51, 49 vacuum, cathode ray, for high temperature, 53, in Sagittarius and Scutum, including Messie 8, 16, and 22, 51, 4 stars in (see Cluster stars) Galaxy Colors (see Cluster stars, Novae, Stars, Varibrightness, surface, as viewed from distance, ables) 52, 162

Galaxy (cont.)

comparison with typical spiral nebulae, 52, 162 distribution of stars in; section, 52, 23

origin; theory, 51, 309

radiation pressure exerted on nebulae, 53, 1 Giant stars (see Cluster stars and Variables) Gratings (see Spectra)

polarizing effect on reflected and transmitted light, 51, 129

theory, Rayleigh-Voigt; evidence confirming, 51, 129

Gravitation

attraction on nebulae due to Galaxy, 53, I Einstein effect; photographic measurement; possible error due to mutual action of adjacent images, 52, 98; 53, 349

Majorana's theory; astronomical consequences, 54, 334

Hydrogen; production by high potential sparks, apparent, 52, 47

Integral, definite

table of values of $\int_{0}^{1} (1-x^{2})^{n+\frac{1}{2}} \cos kx \cdot dx, 53,$

Interferometer, stellar (Michelson's) description

100-inch; accuracy and limits, 51, 263 20-foot; construction, adjustments, and photographs, 53, 249

modification to compensate for atmospheric dispersion, 51, 263; 54, 78

effective wave-length, determination, 51, 263 measurements

a Capella; separation and positions of components, and inclination of orbit, 51, 263 theory and uses; to measure

angles, extremely small, and changes of angle, 51, 257

diameters of planetoids, satellites, and stars, 51, 257; also distribution of luminosity on the disks, 53, 249

dispersion of the atmosphere, 51, 263 parallaxes and relative motions of stars, 51,

relative brightness, position, and separation of double stars (binaries), 51, 257, 263

Jean's contributions to theory of cosmic origins, 51, 309

Kirchhoff's law (see Absorption)

Light (see Absorption, Dispersion, Emission, Polarization, Radiation, Refraction, Spectra)

Limonene

dispersion, rotary, magnetic, and natural, 436-620 μμ, 54, 116

refractive indices, 436-620 $\mu\mu$, 54, 116

Magnetic rotary dispersion (see Dispersion)

Magnitudes (see Binaries, Clusters, Cluster stars, Stars, Variables)

Markings, dark (see Nebulae)

Mass, constancy of; possible influence of one body on the mass of another, 54, 334

Melting-point, of palladium, 51, 244

Meteorites, falling into sun; light to be expected from, 51, 37

Molecular models, Bohr type, of halogen acids, as basis for theory of band structure, 51, 230

Naphthalene, a-monobrom-; dispersion, rotary, magnetic, 436-620 μμ, 54, 45

N.G.C.; corrections to descriptions and positions of various nebulae, 51, 276

Nebulae

attraction, gravitational, by Galaxy, 53, 1 brightness, surface, of nebulous areas

measurement, photographic method, 52, 162 spirals; results for well-known nebulae compared with value for Galaxy, 52, 162

catalogue, descriptive, of 330, 51, 276

changes in N.G.C. 1555 and 2245, 51, 276 constant (unchanged); N.G.C. 995, 1186, 2024,

and 7023, 51, 276 dark clouds or markings; photographs

I.C. II, 5146, Cygnus; unique array, 51, 276 N.G.C. 2146, Camelopard, 51, 276

near & Orionis, including Barnard 33, 53, 392 in Sagittarius and Scutum; N.G.C. 6523 (M8),

6611 (M16), and 6618 (M17), 51, 4

forces acting on; electrostatic, gravitational, and radiative, 53, 1

internal motion in spirals (see Messier 51, 81) I.C. 431, 432, and 434; photographs, 53, 392

I.C. II, 5146, Cygnus; dark markings; photograph, 51, 276

Messier 8 (N.G.C. 6523); photograph, 51, 4 Messier 16 (N.G.C. 6611); photograph and spectrum, 51, 4

Messier 17 (N.G.C. 6618) (Swan); photograph, 51,4

Messier 51 (spiral); internal motion, photograph, and proper motion, 54, 237

Messier 81 (spiral); internal motion, photograph, and proper motion, 54, 347

N.G.C. 1555 and 2245; changes in, 51, 276 N.G.C. 1700 and 3379; radial velocity and

spectrum, 51, 276

N.G.C. 2023 and 2024; photograph, 53, 392

N.G.C. 2146; photograph, 51, 276

N.G.C. 2261 (Hubble), theoretical explanation, 53, 160

N.G.C.; corrections to descriptions and positions of various nebulae, 51, 276

new; descriptive catalogue of 255, 51, 276 origin, theory of

N.G.C. 2261, with fan-shaped appendage, 53, 169

Nebulae (cont.) Photo-electric photometer, for stars origin, theory of (cont.) description, 51, 193; precision, 53, 105; 54, 81 planetary nebulae, by collision, 54, 229 Photographic plates; properties spirals, 54, 347; Jeans' theory, 51, 309 contrast functions γ and Γ ; variation with photographs, large scale (see I.C. 431, 432, 434; wave-length; Purkinje effect, 52, 86 I C. II, 5146; M8, 16, 17, 51, 81; N.G.C. drying, time of; effect of exposure on, 53, 349 2023, 2024, 2146; also dark clouds) grain size; relation to sharpness, theoretical, near & Orionis, 53, 392 52, 201 images in Sagittarius and Scutum, 51, 4 spirals; N.G.C. 2146 and 15 others, 51, 276; contraction and distortion during drying, 52, 98 M51, 54, 237; M81, 54, 347 mutual action of adjacent spectrum lines and misc.: I.C. I, 1470; N.G.C. 1491, 2245, 2247, stars; analysis of effects involved; varia-2294, 2359, 3379, 3384, and 6888, 51, 276 tion with exposure, 53, 349 planetary, origin; collision theory, 54, 229 sections of star images, Plates. 52, 86, 98 proper motion (see Messier 51 and 81) intensification; effect on resolving power, 52, 201 radial velocity penetration of light of different wave-lengths, explanations suggested; discussion, 53, 1 52, 86 N.G.C. 1700 and 3379, 51, 276 resolving power radiation pressure on, due to Galaxy, 53, 1 effect of intensification. 52, 201 spectrum (see M16: N.G.C. 1700 and 3370) relation to grain size and sharpness, theonebulosity around six nebulous stars, 52, 8 retical, 52, 201 spirals (see brightness, catalogue, internal variation with development and wave length motion, new, origin, photographs, proper 52, 201 motion) sharpness Swan (N.G.C. 6618); photograph, 51, 4 measurement; improved method, 52, 201 theory (see N.G.C. 2261, origin, radial velocity) relation to contrast and turbidity, 52, 201 effect of passage of a star near or through a variation with development and wave-length nebula, 53, 169; 54, 229 types, four; examples, 53, 392 shifts of spectrum lines and star images due to variable (see changes) mutual action of adjacent images (see Nebulous areas; brightness (see Nebulae) images) Nebulous stars (see Stars) theoretical relations for resolving power and sharpness, 52, 201 brightness and color of Persei No. 2, 52, 183 turbidity; variation with wave-length, 52, 201 definition suggested, 54, 220 Photographs (see Clusters, Comets, Nebulae, Sun) Ophiuchi 1919; spectrum and radial velocity, Photometer (see Photo-electric photometer) 51, 121 Photometry, photographic, heterochromatic; disorigin, collision theory, 54, 229 cussion and warning, 52, 86 spectrum Planetoids, diameter; measurement, interferom-Ophiuchi 1919; shifts of lines and bands, eter method, 51, 257 51, 121 Polarization shifts, simple interpretation, 51, 121 measurement, using gratings; warning, 51, 129 theory (see origin) produced by gratings, reflected and transmitted light, 51, 129 Occlusion, of RaEm in discharge tubes, 54, 285 Prominences, solar (see Sun) Origin; theory of (see Binaries, Galaxy, Nebulae, Proper motion (see Binaries, Nebulae, Stars) Novae, Solar system, Variables) Purkinje effect, photographic, 52, 86 cosmic origins; Jeans' contributions, 51, 309 Quantum theory (see Spectra) Ozone, in atmosphere (see Atmosphere) Palladium, melting-point, 51, 244 Radial velocity (see Binaries, Nebulae, Novae, Parallax (see Binaries, Clusters, Stars, Variables) Photo-electric cells Radiation alkali metals and hydrides; color sensitiveness constant c2; computed from gold point to curves, 52, 129 palladium point ratio, 51, 244 fatigue tests for K and KH cells, 52, 129 pressure preparation of Li cell, 52, 129 on atoms and electrons; theory based on sensitiveness, color, for 30 cells, including all classical dynamics, 52, 65 alkali metals and hydrides of Na, K, Rb, between binary stars, 53, 1 and Cs, 52, 129 on nebulae due to Galaxy, 53, 1

```
occlusion in discharge tube, 54, 285
                                                            acids, Swan, water, C, CO, CO2, CF4,
  purification; modification of Duane's apparatus,
                                                            H2S, N. SO2)
                                                          arc and furnace spectra; comparison of in-
      54. 285
Refractive indices (see Camphor, Limonene,
                                                            tensity distribution; CN and Swan bands,
                                                            53, 161
      Sugar, Tartaric acid)
                                                          Deslandres' law; test with nitrogen positive
Satellites, diameter; measurement; interferometer
                                                            bands, 52, 301
                                                          theory of structure of infra-red bands (see
      method, 51, 257
                                                            halogen acids)
Scandium, carbide; possible formation in electric
                                                          vacuum tube discharge through COa, HaS,
      furnace, 54, 28
                                                            NH4, N2O, N2O2, and SO2; bands excited
Sky, night; brightness; various determinations;
                                                            in visible and ultra-violet, 52, 301
      discussion, 52, 123
Solar corona and prominences (see Sun)
                                                          absorption spectrum, electric furnace; varia-
Solar system, origin; tidal theory, 51, 309
                                                            tion with temperature, 51, 13
Spectra and spectroscopy
                                                          arc spectrum, low current; variation with
  absorption spectra
                                                            current, 54, 191
    exploded wire, of Fe, \lambda 2270-5645 A, spectro-
                                                          classification of lines
      gram, 51, 37
    furnace, electric, of Ba, Ca, Co, Fe, Ni, and
                                                            furnace excitation, infra-red, 51, 179
                                                            low-current arc excitation, 54, 101
      Ti; variation of relative intensities with
                                                          infra-red furnace spectrum to 856 μμ, at vari-
      temperature, 51, 13
                                                            ous temperatures, 51, 179
    production of
                                                          series of single lines and triplets; identifica-
      electric furnace spectra, 51, 13
                                                            tion of terms; constants, 51, 23
      high temperature spectra, extreme; ex-
                                                        binaries (see Binaries)
        ploded wire method, 51, 37
                                                        hismuth
    theory (see halogen acid gases)
                                                          arc spectrum, low current; variation with
      Kirchhoff's law applied to electric furnace
                                                            current, 54, 246
        spectra, 51, 13
                                                          classification of lines, arc excitation, 54, 246
    Zeeman effect, inverse (see Fe and V)
                                                          structure
  air
                                                            of \u03bb \u03bb 4122, 4308, 4722; Plate, 53, 323, 339
    arc spectrum; two new lines, 54, 246
                                                            of \(\lambda\) 3397, 3511, 3596; Plate only, 53, 323
    spark spectrum, condensed
                                                            of $\lambda 4722; changes in relative intensity of
      effect of self-inductance on relative intensi-
                                                              components, 53, 339
        ties, 590-872 μμ, 51, 236
      identification of Ar, N, and O lines, 500-
                                                          arc spectrum, low current; variation with
        872 μμ, 51, 236; 54, 76
                                                            current, 54, 246
      shift with reference to vacuum tube lines,
                                                          classification of lines, arc excitation, 54, 246
                                                          new resonance line, \(\lambda\) 3779, 54, 246
      wave-lengths, 590-872 \mu\mu, 51, 236; 54, 76
                                                        calcium
  ammonia bands, visible and ultra-violet;
                                                          absorption spectrum, electric furnace, 51, 13
      identification, 52, 301
                                                          arc spectrum, low current; variation with
  arc spectra (see air, Fe, pole-effect, pressure shift)
                                                            current, 54, 191
    anode and cathode spectra; relative behavior
                                                          classification of lines
      of various metal lines; variation with
                                                            furnace excitation, infra-red, 51, 179
      atomic weight, 54, 65
                                                            low-current arc excitation, 54, 191
    comparison with furnace spectra (see Ca,
                                                          furnace spectra; absorption, emission and
      cyanogen, and Swan bands)
                                                            mixed, 51, 13; infra-red, 51, 179
    ionization lines; behavior, 54, 191, 246
                                                          infra-red furnace spectrum to 733 \mu\mu, at
    low-current; variation with current; excita-
                                                            various temperatures. 51, 179
      tion stages
                                                          new lines, fifty, 52, 265
        for Ba, Ca, K, Mg, Na, Sr, 54, 191
                                                          pressure shifts to 1 atm., 315-650 μμ, 53, 224
        for Ag, Bi, Cd, Cu, and Zn, 54, 246
                                                          series of singlets and triplets; identification
    relation of results to Bohr theory and Ritz
                                                            of terms; constants, 52, 265
      equations, 54, 246
    standard lines, secondary; comparison of
                                                          bands, negative; origin, 52, 301
      12-mm 5-amp. with 6-mm 6-amp. arc,
      53, 260
                                                          band, positive, fourth; wave-lengths, 52, 301
  argon; lines in condensed spark spectrum in air,
                                                          ultra-violet vacuum spark spectrum; à 1931-
      590-872 μμ, 51, 236; 54, 76; shift with
                                                            360 A; Plate, 52, 47; wave-lengths, 53, 150
      reference to vacuum tube lines, 51, 236
                                                          X-ray spectrum, L-series; identification, 52, 47
```

band spectra (see ammonia, cyanogen, halogen

Radium emanation (Niton)

comparison with arc spectra Spectra (cont.) carbon dioxide; bands in visible and ultraintensity distribution in violet, including several new, 52, 301 cyanogen and Swan bands, 53, 161 carbon monoxide; bands in visible and ultraspectra of Ba, Ca, Co, Ni, Sr, 51, 170; violet; identification, 52, 301 Sc, 54, 28 carbon tetrafluoride; bands in visible and comparison with solar spectrum; Sc, 54, 28 ultra-violet; wave-lengths, 54, 133 effect of small potential gradient, 52, 187 cathode rays; intensity of N bands excited, as infra-red absorption spectra to 920 μμ, of function of energy, 52, 278 Ba, Ca, Co, Ni, and Sr, 51, 170 classification of lines mixed absorption and emission spectra; enhanced lines, Fowler's; discussion, 54, 246 production, method, 51, 13 furnace excitation stages, King's electric; origin of radiation; discussion, 52, 187 for Ba, Ca, Co, Ni, and Sr; infra-red, 51, 179 red fringe; explanation, 52, 187 for Mn, λ 2795-6500 A, 53, 133 variation with temperature (see furnace absorpfor Sc, λ 3015-6559 A, 54, 29 tion spectra) low-current arc excitation stages Zeeman effect for iron lines, 51, 107 for Ag, Bi, Cd, Cu, and Zn, 54, 246 grating spectrograph for Ba, Ca, K, Mg, Na, and Sr, 54, 191 comparison with interferometer, 53, 260 comparison with furnace classification, 54, ghosts and reversals; use of, in accurate 246 measurements, 53, 260 intensity effect on wave-length nil, 53, 260 absorption spectrum, electric furnace; variapolarizing effect on reflected and transmitted tion with temperature, 51, 13 light, 51, 129 classification of lines; furnace stages; infraultra-violet, extreme, 52, 47, 286; 53, 150 red to 809 $\mu\mu$, 51, 179 infra-red furnace spectrum to 809 μμ, at halogen acid gases (HBr, HCl, HF); bands various temperatures, 51, 179 HCl band 3.7 μ; wave-lengths, law of spacing, evidence of satellites, 53, 300 arc spectrum, low current; variation with theory of structure of infra-red bands current, 54, 246 isotopic theory of doublets, 52, 248 classification of lines; low-current arc stages quantum theory, based on simple molecular and comparison with furnace stages, 54, 246 model of Bohr type, 51, 230 continuous background obtained with exhelium ploded wire source, 51, 37 excitation of various spectra; minimum cyanogen bands voltage, 52, 1 intensity distribution, in furnace and arc intensity, relative, of series lines and bands in spectra, 53, 161 arc spectrum; variation with voltage, 52, 1 λ 3883 in arc and furnace; Plate, 53, 161 ultra-violet, extreme, spark spectrum; identiexcitation of spectra; minimum voltage; hefication of lines, 52, 47 lium spectra, 52, 1; variation with current density; helium, 52, 1 hydrogen, Balmer series shift of Ha; condensed spark in air comexploded wire spectra (extremely high temperapared with vacuum spectrum, 51, 236 ture absorption and emission spectra) variations in relative intensity of lines in of Cu, Ni, Mn; continuous background, 51, 37 spectrum of Class Md variable star, 53, 185 of Fe; absorption spectrum, λ 2270-5645 A; Plate, 51, 37 hydrogen bromide, chloride, and fluoride (see production, method, 51, 37 halogen acid gases) use for study of pressure shift suggested, 51, 37 hydrogen sulfide; spectrum of discharge through fluorescence spectra (see mercury, 54, 149) 52, 301 fluorine; spark spectrum, visible and ultrainfra-red spectra (see air, Ba, Ca, Co, Ni, Sr, violet, of pure gas; wave-lengths, 54, 133 halogen acid gases, sun, water) Fraunhofer lines (see Sun, spectrum) elimination of scattered light of shorter furnace spectra, electric wave-lengths, 53, 121 screen for light to 7200 A, 53, 121 absorption spectra of metallic vapors comparison with emission spectra, 51, 13 interferometer spectrograph production, method, 51, 13 comparison with grating, 53, 260 theory; Kirchhoff's law applied to, 51, 13 ghosts and reversals; use of, for accurate variation with temperature; spectra of measurements, 53, 260 Ba, Ca, Co, Fe, Ni, and Ti, 51, 13 intensity effect on wave-length nil, 53, 260 Ba, Ca, Co, Ni, Sr; infra-red, 51, 179 reduction of measurements, method, 53, 260 Mn, 280-820 μμ, **53,** 133 ionization lines, in low-current arc; variation Sc, 301-656 µµ, 54, 28 with current, 54, 191

52, 47; wave-lengths, 53, 150 iron absorption spectrum nitrogen; band spectrum furnace; variation with temperature, 51, 13 excitation by vacuum discharge through exploded wire, \(\lambda\) 2270-5645 A; Plate, 51, 37 N2O and N2O2, 52, 301 arc lines, \(\lambda\) 3370-6750 A; wave-lengths of intensity as function of energy of the exciting 1026 lines, measured with grating and cathode rays, 52, 278 interferometer, and compared with Bureau new ultra-violet positive bands from lowof Standards results, 53, 260 current arc in air, possible, 54, 246 intensity effect for arc lines nil, 53, 260 structure; divergence from Deslandres' law, pole-effect in Pfund arc 52, 301 comparison of 12-mm 5-amp, with 6-mm wave-lengths of third positive, 52, 301 6-amp. arc, 53, 260 nitrogen; line spectrum relation to Zeeman effect nil, 53, 329 identification of lines in condensed spark in ultra-violet spark spectrum to 200 A; Plate, air, 590-870 μμ. 51, 236; 54, 76 52, 47; wave-lengths, 53, 150 relative intensity; effect of self-inductance variations in relative intensity of lines in with spark, 51, 236 spectrum of Class Md variable stars, 53, 185 shift of spark lines with reference to vacuum Zeeman effect tube lines, 51, 236 furnace lines: direct and inverse effect for nitrogen peroxide; spectrum of discharge 100 lines; camparison with results for through, 52, 301 spark lines; Plates, 51, 107 nitrous oxide; spectrum of discharge through, relation to pole-effect nil, 53, 329 52, 301 isotopes, components due to novae; spectrum (see Novae) measurement of separation; displaceable oxygen slit-method suggested, 53, 329 identification of lines in spark in air and O2, theory, for case of bands of HBr, HCl, 52, 248 590-870 μμ, 51, 236; 54, 76 shift of spark lines with reference to vacuum magnesium tube lines, 51, 236 arc spectrum, low-current; variation with pole-effect current, 54, 191 classification of lines; low-current arc stages, in iron arc, comparison of 6-mm 6-amp. and 12-mm 5-amp. arcs, 53, 260 54, 191 measurement, displaceable slit-method sugmanganese gested, 53, 320 classification of lines, 280-650 μμ; furnace relation to Zeeman effect, for iron, nil, 53, 329 stages, 53, 133 potassium furnace spectrum, 280-820 μμ; variation with temperature, 53, 133 arc spectrum, low-current; variation with measurements current, 54, 191 classification of lines, low-current arc stages, errors, possible, due to mutual influence of adjacent photographic images, 52, 98; 54, 191 pressure shift 53, 349 calcium arc lines, 315-650 µµ, 53, 224 of shifts, minute; displaceable slit method. source for study; exploded wire suggested, suggested, 53, 329 of ultra-violet wave-lengths, extremely short, in stellar spectra, Arcturus, Procyon, 53, 327 53, 150 quantum theory (see theory) mercury; fluorescence spectrum, excitation of; active molecules; relation to exciting radium emanation (niton) spectrum, 54, 149 new lines, $398-745 \mu\mu$, **54,** 285 relative intensity of lines; variation during mixed absorption and emission spectra; production in electric furnace, 51, 13 discharge, 54, 285 resonance lines in low-current arc spectra nebulae; spectra (see Nebulae) neon; low-voltage spectrum of trace of neon in intensity variation with current for helium, 52, 1 Ag, Bi, Cd, Cu, and Zn, 54, 246 Ba, Ca, K, Mg, Na, and Sr, 54, 191 Ritz equations (see series) absorption spectrum, electric furnace; variation with temperature, 51, 13 scandium classification of lines, furnace stages; infraclassification of lines, furnace stages, 54, 28 red to 780 μμ, 51, 179 furnace spectrum, 301-656 μμ; at various infra-red furnace spectrum to 920 $\mu\mu$; at temperatures; comparison with arc, solar various temperatures, 51, 179 and sun-spot spectra, 54, 28

ultra-violet spark spectrum to 731 A; Plate,

Spectra (cont.)

theory Bohr; relation of low-current arc results to, series in Ba spectrum, singlets and triplets, 51, 23 54, 246 relation of minimum voltage results for in Ca spectrum, singlets and triplets, 52, 265 helium to, 52, 1 notation; explanation, 51, 23 isotopic, of separation of doublets of HBr and Ritz equations; relation of low-current arc HCl, 52, 248 quantum, of structure of band spectra of results to, 54, 246 shifts of lines (see spark spectra, Sun) halogens, based on simple molecular model measurement; displaceable slit method sugof Bohr type, 51, 230 gested, 53, 329 titanium; absorption spectrum, electric furnace, photographic, due to mutual action of adjacent at various temperatures, 51, 13 images, 52, 98; 53, 349 ultra-violet spectra (see Ca, CO, CO2, F, Mn, N, silver NH4, N2O, N2O2, SO2, Sun) arc spectrum, low current; variation with extreme, to 200 A (see C, He, Fe, Ni, Zn) current, 54, 246 measurements of wave-lengths, 53, 150 classification of lines; low-current arc stages, source; condensed vacuum spark, 52, 286 54, 246 spectrograph, vacuum, 52, 47, 286; 53, 150 sodium screens for region 290-315 μμ, 54, 297 arc spectrum, low current; variation with spectrograph, special double, for solar speccurrent, 54, 101 trum, 290-315 μμ, 54, 297 classification of lines; low-current arc stages vacuum discharge spectra (see CO, CO2, CF4, 54, 191 N, Ne, NH4, RaEm, SO2) Zeeman effect for D-lines; explanation of intensity of N bands as function of energy of Woltjer's observations, 51, 107 cathode rays, 52, 278 sources of light (see arc, exploded wire, spark, minimum voltage for excitation of He spectra, 52, I exploded wire; appearance and mechanical spectra of discharge through CO2, H2S, NH4, N2O, N2O2, SO2, 52, 301 effects of explosion, 51, 37 spark spectra (see air, argon, fluorine, hydrovacuum sources gen, nitrogen, oxygen) cathode rays from treated carbon cathode, relative intensity; effect of self-inductance on used to heat anode, 53, 323 air and O lines, 51, 236 condensed spark, 52, 286 shift with reference to vacuum tube lines; air, vanadium; Zeeman effect, direct and inverse, Ar, H, N, and O lines, 51, 236 for 90 furnace lines; comparison with effect for spark lines, 51, 107 ultra-violet, extreme (see C, Fe, He, Ni, Zn, Venus; spectrum (see Venus) ultra-violet) water vapor; absorption band, 930-963 μμ, spark, condensed, in vacuum; as source for extreme ultra-violet, 52, 286 53, 121 spectrographs (see grating, interferometer) X-ray spectra standards, international secondary; questioned L-series of carbon; identification, 52, 47 lines; comparison of 6-mm 6-amp, and Zeeman effect (see Fe, Na, Sc, V) 12-mm 5-amp. arcs; pole-effect, 53, 260 furnace lines compared with spark lines, of Fe and V, 51, 107 stellar spectra (see Stars) strontium relation to pole-effect for Fe lines nil, 53, 329 arc-spectrum, low-current; variation with current, 54, 191 arc spectrum, low current; variation with classification of lines current, 54, 246 arc stages, low-current, 54, 191 classification of lines, low-current arc stages, furnace stages, infra-red, 51, 179 54, 246 infra-red furnace spectrum to 920 μμ; at ultra-violet spark spectrum to 316 A: Plates, various temperatures, 51, 179 52, 47, 286; wave-lengths, 52, 286 structure (see bands, bismuth) Stars sulfur dioxide bands atmospheres, pressure in; Arcturus and Procyon, in visible and ultra-violet, including forty new; wave-lengths, 52, 301 binaries (see Binaries and Variables) sun (see Sun) brightness (see parallax)

Swan band

intensity distribution; 5165 A; arc and

furnace spectra, 53, 161

Spectra (cont.)

scandium (cont.)

Zeeman effect for lines in sun-spot spectrum,

catalogue (see Cluster stars)	323
in Pleiades; magnitudes and colors of 82r	scales, photo-visual; comparison of Barnard's
stars in region 2° square, 54, 323	with Mount Wilson, 54, 323
spectroscopic parallaxes, magnitudes, type,	measurements from photographs; possible errors
and proper motion of 1646 stars, 53, 13	due to contraction effect, 52, 98; and
Cepheids (see Variables)	mutual action of images. 53, 349
Class Bo-B5; statistical study of 180 stars;	nebulous stars
mean magnitude, parallax and proper	colors of 47, including ρ Ophiuchi, σ and 22
motion, 54, 140	Scorpii, 52, 8
Class O; collision theory of origin, 54, 229	spectrum
Class Oe5 (see Binaries)	R Aquarii, nebulous lines; Plate, 53, 375
Class Md (see Variables)	nebulosity around six stars, 52, 8
cluster (see Cluster stars)	novae (see Novae)
colors (see Cluster stars)	· ·
determination for nebulous stars, 52, 8	number of each absolute magnitude; luminosity
	curve, 52, 23 (see density)
nebulous stars, 47, including ρ Ophiuchi, σ	parallaxes and absolute magnitudes
and 22 Scorpii, 52, 8	catalogue for 1646 stars; spectroscopic and
in Pleiades; 753 dwarfs, 54, 323	trigonometric results, 53, 13
comparison stars, for Messier 72; photographic	Class Bo-B5 stars; mean for 180, 54, 140
magnitudes of 29, 52, 232	Boss 1517, 5r , 254
constant stars, from photometric studies	determination; methods
Bond 624, 53 , 317	accuracy, relative; discussion, 54, 140
r and σ Cassiopeiae, 54, 8r	interferometer method suggested, 5r, 257
Class B: π^4 Orionis, ξ , e, and μ Tauri, 51, 193	spectroscopic; description, 53, 13
π4 Orionis, 51, 193 , 218	distribution of stars in space and number of
ξ, e, and μ Tauri, 5r, 193	each magnitude, 52, 23
l and π Persei, 53 , 105	relation of magnitude to space velocity;
radial velocity constant (see radial)	statistical study of 1350 stars, 54, 9
density of stars	relation of parallax to apparent magnitude
in Gal. long. $+32^{\circ}$, lat. -20° , 52, 73	and proper motion, 52, 23
	pressure in atmospheres (see atmospheres)
in Galaxy; distribution, 52, 23	pressure in atmospheres (see atmospheres)
in Galaxy; distribution, 52, 23 in space, as function of parallax and magni-	proper motion
in Galaxy; distribution, 52, 23 in space, as function of parallax and magni- tude, 52, 23	proper motion catalogue, for 1646 stars, 53, 13
in Galaxy; distribution, 52, 23 in space, as function of parallax and magni- tude, 52, 23 diameters	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B ₅ , mean for 180 stars, 54, 140
in Galaxy; distribution, 52, 23 in space, as function of parallax and magni- tude, 52, 23 diameters measurement with stellar interferometer, 51,	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237
in Galaxy; distribution, 52, 23 in space, as function of parallax and magni- tude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method sug-
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-Bs, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space)
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-Bs, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries)
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2395, +83°104 Cephei, 52, 198
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2395, +83°104 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi,
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2395, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume,	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 10, +59°2395, +83°104 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 10, +59°2395, +83°104 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +50°2305, +83°104 Cephei, 52, 108 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +50°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars)
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2395, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, r98 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, r98 spectra (see nebulous stars) combined bright and dark line spectra;
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for 180 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2395, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 a Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables)
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) a Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax)	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables)
in Galaxy; distribution, 52, 23 in space, as function of parallax and magnitude, 52, 23 diameters measurement with stellar interferometer, 51, 257 (see Interferometer) α Orionis, 53, 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53, 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52, 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei,	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) a Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105 magnitudes, photographic and photo-visual	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +50°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) a Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, r98 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, r98 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13 theory; effect of passage of a star near or
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) a Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105 magnitudes, photographic and photo-visual catalogue, of 1646 stars, 53 , 13	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 51, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 10, +59°2395, +83°104 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13 theory; effect of passage of a star near or through a nebula, 53, 169; 54, 229
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) α Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105 magnitudes, photographic and photo-visual catalogue, of 1646 stars, 53 , 13 Class Bo-B5, 180 stars, 54 , 140 comparison stars	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, r98 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, r98 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13 theory; effect of passage of a star near or through a nebula, 53, 169; 54, 229 origin of Class O stars, 54, 220 variable stars (see Variables)
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) a Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Gulster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105 magnitudes, photographic and photo-visual catalogue, of 1646 stars, 53 , 13 Class Bo-B5, 180 stars, 54 , 140 comparison stars for Messier 72, 29, photographic, 52 , 232	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +50°2395, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13 theory; effect of passage of a star near or through a nebula, 53, 169; 54, 229 origin of Class O stars, 54, 229 variable stars (see Variables) velocity; radial, tangential and space
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) a Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Cluster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105 magnitudes, photographic and photo-visual catalogue, of 1646 stars, 53 , 13 Class Bo-B5, 180 stars, 54 , 140 comparison stars for Messier 72, 20, photographic, 52 , 23 2 for Nova Persei No. 2, 36, 52 , 183	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +59°2305, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, r98 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, r98 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13 theory; effect of passage of a star near or through a nebula, 53, 160: 54, 229 origin of Class O stars, 54, 229 variable stars (see Variables) velocity; radial, tangential and space distribution; frequency of each velocity;
in Galaxy; distribution, 52 , 23 in space, as function of parallax and magnitude, 52 , 23 diameters measurement with stellar interferometer, 51 , 257 (see Interferometer) a Orionis, 53 , 249 disk (see diameter) darkening of limb (see Variables) distribution of luminosity; interferometer method of study, 53 , 249 distribution in space (see density) number of each magnitude per unit volume, luminosity curve, 52 , 23 double stars (see Binaries, visible) dwarf stars, in Pleiades (see catalogue) Galaxy (see Galaxy) giants (see Gulster stars, Variables) magnitudes, absolute (see parallax) magnitudes, photo-electric, of β, δ, l, π Persei, 53 , 105 magnitudes, photographic and photo-visual catalogue, of 1646 stars, 53 , 13 Class Bo-B5, 180 stars, 54 , 140 comparison stars for Messier 72, 29, photographic, 52 , 232	proper motion catalogue, for 1646 stars, 53, 13 Class Bo-B5, mean for r80 stars, 54, 140 large; two faint stars near M51, 54, 237 measurement; interferometer method suggested, 57, 257 radial velocity (see velocity, space) binaries (see Binaries) 16, 19, +50°2395, +83°r04 Cephei, 52, 198 constant for δ Ophiuchi, η Serpentis, β Pegasi, 52, 317 +73°835 Draconis, 52, 198 α Hercules and comp., 53, 201 25 θ Ursae Majoris, 23 δ Ursae Minoris, 52, 198 spectra (see nebulous stars) combined bright and dark line spectra; explanation, 51, 13 variable (see Variables) spectral type catalogue for 1646 stars, 53, 13 theory; effect of passage of a star near or through a nebula, 53, 169; 54, 229 origin of Class O stars, 54, 229 variable stars (see Variables) velocity; radial, tangential and space

in Pleiades; statistical study of 82r, 54,

Stars (cont.)

Sugar, cane; aqueous solution

dispersion, rotary, magnetic and natural, 436~ 620 μμ, 54, 116

refractive index, one to one solution, 436-620 μμ, **54,** 116

corona; May 29, 1919; photograph, 51, 1 eclipse; May 29, 1919; Smithsonian expedition; brief report, 51, 1 (see corona)

meteorites falling into; light expected from; theoretical discussion, 51, 37

prominence; May 29, 1919; photograph, 51, 1 October 8, 1920; very high; stages of growth; photographs with Ca line, 53, 310

spectrum, including Fraunhofer lines

comparison with arc and furnace spectra; scandium lines, 54, 28

energy distribution, corrected for atmospheric absorption, 315-290 μμ, 54, 297

infra-red, 890-990 μμ; identification and origin of lines; wave-lengths of 563 lines, including 50 solar; also Plate, 53, 121 shifts of lines

atmospheric refraction, 53, 380

center-aic, predicted by anomalous dispersion theory, 54, 92

mutual influence of adjacent lines; data from limb-center comparisons, 54, 92 (see photographic)

photographic effect of adjacent lines, 52, 98; 53, 349

theory, anomalous dispersion; of gravitational shift, limb-center shift, and mutual influence shift, 54, 92

spot spectra (see spots)

theory (see shifts)

ultra-violet (see energy distribution)

map, photographic; 315-290 μμ, 54, 297 spots

spectrum

comparison with arc and furnace spectra; scandium lines, 54, 28

Zeeman effect for scandium lines, 54, 28 theory; cooling of rising gases, 54, 293 theory (see spectrum, spots)

Tartaric acid; aqueous solution

dispersion, rotary, natural, and magnetic, 436-620 μμ, **54,** 116

refractive index, for one to one solution, 436-620 μμ, **54,** 116

Telescope objective; diffraction by; effect on image of disk and combination of disks, including lune; mathematical theory, 51, 73

Temperature scale

brightness ratio, gold point to palladium point, 51, 244

palladium melting-point, 51, 244

Theory (see Binaries, Dynamics, Electron, Gratings, Gravitation, Interferometer, Nebulae, Novae, Origin, Photographic plates, Radiation pressure, Spectra, Stars, Sun, Variables

Transits; observation; effect of diffraction by telescope objective; theory, 51, 73

Triple systems

65 τ Cygni; orbit of spectroscopic binary component, 53, 144

13 Ceti; orbit of spectroscopic binary component and perturbations due to fainter visual component, 52, 110

light curve of variable binary component of λ Tauri, 51, 193

orbits; of binary components (see 65 τ Cygni, 13 Ceti, and \(\text{Tauri} \)

perturbations due to third body (see 13 Ceti and \(\lambda\) Tauri)

λ Tauri; photometric study of variable binary component; orbit, light curve, effect of third body nil, 51, 193

Variable nebulae (see Nebulae)

Variable stars

Algol; photometric study, elements of eclipsing system, light curve, color of satellite, 53,

R Aquarii; spectrum, intensities and displacements of lines and nebular lines (Plate), 53,

γ Argus; spectrum; temporary, shifting, absorption He lines, 52, 39

γ Camelopardalis; photometric study; light curve and elements, 54, 217

RS Canum Venaticorum; light curve, elements, computed parallax, 53, 99

l Carinae; periodic variations of wave-length and spectral type, 54, 161

H. Cassiopeiae; photometric study; light curve, elements, darkening of limb, 54, 81 SX Cassiopeiae; light curve, elements, 53, 165

T Cephei; periodic spectrum changes, 53, 185

U Cephei; photometric study; light curve, elements, evidence of tidal evolution, 52,

Cepheids (see l Carinae, X Cygni, Messier 72) light curves of 26, in Messier 72, 52, 232

light range, small; possibility of Cepheids with; suggestion, 51, 62

orbit, spectroscopic, for X Cygni. 53, 95 origin; collision theory of, 54, 229

magnitude, mean, for 26 in M72, 52, 232 periods; in M72, 26 variables, 52, 232 relation to spectral type, 54, 161

spectral type; range of variation and relation to period, 54, 16

spectrum; periodic variations in wavelength and type; I Carinae, 54, 161 theory (see origin)

general conclusions; mean atomic weight; ratio of mass to radius, 52, 73 relation of period to brightness, 52, 73

variation

binary theory; discussion, 51, 62 condition of; ratio mass to radius, 52, 73 Variable stars (cont.) Class Md; spectrum; periodic changes in emission lines (see T Cephei, X Cygni, R Hydrae, R Leonis, X Ophiuchi, R Serpentis, 53, 185; R Aquarii, 53, 375) classification of long-period variables, 53, 179 cluster variables (see Cepheids) color variation of typical, 51, 49 colors and magnitudes of 28 in M68, 51, 49 magnitude, absolute, of typical, 51, 49 new; in M3; 17 probable, 51, 140 in M53, 56, 72, and 75; positions of 80 shown on photographs, 52, 73 in M68; 28, mostly typical, 51, 49 colors (see cluster) satellite of Algol, 53, 105 X Cygni; spectroscopic orbit, 53, 95 χ Cygni; periodic changes of spectrum, 53, 185 darkening of limb; 1 H. Cassiopeiae, 54, 81 205 Draconis, probable eclipsing variable, 53, 201 (see Binaries) eclipsing variables light curve and elements (see Algol, y Camelopardalis, RS Canum Venaticorum, SX and 1 H. Cassiopeiae, U Cephei, RT Lacertae, \(\lambda\) Tauri) probable, 205 Draconis, 53, 201 ellipsoidal variable (see π5 Orionis) evolution, tidal, of U Cephei; evidence, 52, 145 giants (see Cepheids) R Hydrae; variations in spectrum, 53, 185 irregular variables; collision theory of origin, 54, 229 RT Lacertae; light curve and elements, 52,

R Leonis; variations in spectrum, 53, 185 light curves (see Cepheids, eclipsing, ellipsoidal, and long-period variables) long-period variables classification, 53, 169 light curves of 66; constants, 53, 169 spectrum of R Aquarii, 53, 375 magnitude (see Cepheids) magnitude, absolute (see cluster, parallax) Messier 3, 53, 56, 68, 72, 75; new variables (see cluster) Messier 72; light curves of 26 Cepheids, 52, 232 nebulous variable (see R Aquarii) new variables (see cluster variables) B.D. +81°27; +81°30, 52, 145 orbits; elements (see eclipsing variables) spectroscopic; X Cygni, 53, 95 origin, of Cepheids and irregular variables; collision theory, 54, 229 X Ophiuchi; variations in spectrum, 53, 185 π5 Orionis; photometric study; light curve and elements, 51, 218 parallax and absolute magnitude, of RS Canum Venaticorum, 53, 99 periods (see light curves, Cepheids) photometric study (see Algol, \(\gamma \) Camelopard, 1 H. Cassiopeiae, U Cephei, λ Tauri) π4 Orionis, constant star, 51, 193, 218 R Serpentis; variations in spectrum, 53, 185 spectrum; periodic variations in emission lines (see R Aquarii, y Argus, Cepheids, Class Md) suspected variables B.D. +10°1771, 52, 9 Bond 624 in Orion; photometric study; Hartwig's elements incorrect, 53, 317 205 Draconis, 53, 201 δ Persei, 53, 105 λ Tauri; photometric study; light curve and elements for binary, 51, 193 theory (see origin)

theory (see origin)
tidal evolution; evidence of; U Cephei, 52, 145
Wolf-Rayet star (see γ Argus)
Venus; spectrum; systematic shifts of solar lines;
variation with zenith distance; explanation;
Plate, 53, 380

X-rays

L-series of carbon; identification, 52, 47 source; condensed vacuum spark, 52, 47

Zeeman effect (see Spectra)



